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Abstract. Modeling propagation of VLF electromagnetic beams in the waveguide earth-ionosphere (WGEI) is of a great importance because variation in the characteristics of these waves is an effective instrument for diagnostics the influences on the ionosphere “from above” (Sun-Solar Wind-Magnetosphere-Ionosphere), “from below” (the most powerful meteorological, seismogenic and other sources in the lower atmosphere and lithosphere/Earth, such as hurricanes, earthquakes, tsunamis etc.), from inside the ionosphere (strong thunderstorms and lightning discharges) and even from the far space (such as gamma-flashes, cosmic rays etc.). Thus, VLF became one of the most universal instrument for monitoring the Space Weather in the direct sense of this term, i.e. the state of the Sun-Earth space and the ionosphere as it is particularly determined by all possible relatively powerful sources, wherever they are placed. This paper is devoted mostly to modelling VLF electromagnetic beam propagation in the WGEI. We present a new tensor impedance method for modelling propagation of electromagnetic beams (TIMEB) in a multi-layered/inhomogeneous waveguide. Suppose that such a waveguide, i.e. WGEI, possesses the gyrotropy and inhomogeneity with a thick cover layer placed above the waveguide. Note a very useful and attractive feature of the proposed TIMEB method: in spite of a large thickness of the waveguide cover layer, the proposed effective impedance approach reflects an impact of such a cover on the electromagnetic (EM) waves, which propagate in the waveguide. This impedance approach can be applied for EM waves/beams in layered gyrotropic/anisotropic active media in very wide frequency range, from VLF to optics. Moreover, this approach can be applied to calculations of EM waves/beams propagation in the media of an artificial origin such as metamaterial microwave
or optical waveguides. The results of the modelling the propagation of VLF beams in the WGEI are included. The qualitative comparison between the theory and experimental observation of increasing losses of VLF waves in the WGEI is discussed. The new proposed method and its further development allows the comparison with the results of the future rocket experiment. This method allows to model (i) excitation of the VLF modes in the WGEI and their excitation by the typical VLF sources, such as radio wave transmitters and lightning discharges and (ii) leakage of VLF waves/beams into the upper ionosphere/magnetosphere.

**Keywords** — ionosphere, atmosphere, VLF, tensor impedance, gyrotropy, layered waveguide, beam, electromagnetic wave, boundary conditions, ionospheric disturbances, vertical coupling processes

### 1 Introduction

This paper is dedicated to the propagation in the system Lithosphere–Atmosphere–Ionosphere–Magnetosphere (LAIM) of electromagnetic (EM) waves /beams in the radio range, with particular applications to very low frequencies (VLF). This topic became very actual due to the following reasons. (1) Variation in the characteristics of these waves is now an effective instrument for the diagnostics of “ionospheric weather” as a part of the Space Weather (Hapgood 2017; Yigit et al. 2016; Richmond 1996) in its direct meaning: the state of the Sun-Earth space and the ionosphere in particular determined by all possible sufficiently powerful sources, wherever they are placed. Change in the characteristics (amplitude and phase) of the VLF waves propagating in the waveguide earth-ionosphere (WGEI) reflects the corresponding variations in the ionospheric electrodynamics characteristics (complex dielectric permittivity) and respectively, the influences on the ionosphere “from above” (Sun-Solar Wind-Magnetosphere Ionosphere (WINDMII) (Patra et al., 2011; Koskinen, 2011; Boudjada et al., 2012; Wu et al., 2016), “from below” (the most powerful meteorological, seismogenic and other sources in the lower atmosphere and lithosphere/Earth, such as cyclones and hurricanes (Nina et al., 2017; Rozhnoi et al., 2014; Chou et al., 2015), earthquakes (Hayakawa, 2015; Surkov and Hayakawa, 2014; Sanchez-Dulcet et al., 2015), tsunamis etc. or from inside the ionosphere (strong thunderstorms and lightning discharges, terrestrial gamma-ray flashes or sprite streamers (Cummer et al., 1998; Qin et al., 2012; Dwyer 2012; Dwyer and Uman, 2014; Cummer et al., 2014; Mezentsev et al., 2018). Note that the VLF signals are very important for the merging of the atmospheric physics and space plasma physics with the astrophysics and high energy physics. The corresponding “intersection area” for these two disciplines includes cosmic rays and very popular now objects of investigation – high-altitude discharges (sprites), anomalous X-ray bursts, powerful gamma-ray bursts etc. The key phenomena for the occurrence of all of these objects are runaway electrons with runaway breakdown, and one of the necessary conditions of them is the presence of cosmic rays, consequently these phenomena are intensified during the air showers generating by cosmic particles (Gurevich and Zubin 2001; Gurevich et al. 2009). The runaway breakdown and lightning discharges including high-latitude ones case radio emission both in HF range, which could be observed using LOFAR and other radio telescopes (Buitink et al., 2014; Scholten et al., 2017; Hare 2018), and in the VLF range. Corresponding experimental research include measurements of the VLF characteristics by the international measurement system of the pairs “transmitted-receiver” separated by a distance of a couple of thousand km (Biagi et al.,
Another international system is based on the measurements of VLF characteristics for the characterization of the thunderstorms with the lightning discharges/World Wide Lightning Location Network (WWLLN) (Lu et al. 2019). (2) Intensification of the magnetospheric research, wave processes, particle distribution and wave-particle interaction in the magnetosphere including radiation belts leads to the great interest to the VLF plasma waves, in particular whistlers (Artemyev et al., 2013; Agapitov et al., 2014; Agapitov et al., 2018).

The differences of our proposed model from the known ones used for the simulation of the VLF waves in the WGEI are the following. (1) In distinction to the impedance invariant imbedding model (Shalashov and Gospodchikov, 2010; Kim and Kim, 2016), our model provides optimal, for definite class of problems, which we consider, balance between the analytical and numerical approaches and is, in fact combined analytical-numerical one, basing on matrix sweep, method (Samarskii, 2001). As a result, this model allows to obtain analytically the tensor impedance and at the same time, provides high effectiveness and stability of the modelling. (2) In distinction to the full-wave finite difference time domain (FDTD) models such as (Chevalier and Inan, 2006; Marshall et al., 2017; Yaxin et al., 2012; Azadifar et al., 2017), our method provides very physically clear lower and upper boundary conditions, in particular physically justified upper boundary conditions corresponding to the radiation of the waves propagation in the WGEI to the upper ionosphere/magnetosphere. This allows in a perspective to determine the leakage modes and to interpret not only ground-based, but also satellite measurements of the VLF beam characteristics. (3) In distinction to the models (Kuzichev and Shklyar, 2010; Kuzichev et al., 2018; Lehtinen and Inan, 2009; Lehtinen and Inan, 2008) based on the mode presentations and made in the frequency domain, we use the combined approach, This approach includes condition of the radiation at the altitudes of the F region, equivalent impedance conditions in the lower E region and at the lower boundary of the WGEI, mode approach, and finally, beam method. This combined approach, finally, creates the possibility to interpret adequately data of both ground and satellite detection on the EM wave/beam propagating in the WGEI and these which leakage from the WGEI into the upper ionosphere/magnetosphere. Some other details on the distinctions from the previously published models are given below in Sect. 3.

The methods of effective boundary conditions, in particular effective impedance conditions (Tretyakov, 2003; Senior and Volakis, 1995; Kurushin and Nefedov, 1983) are well-known and can be used, in particular, for the layered metal-dielectric, metamaterial and gyrotropic active layered and waveguiding media of different types (Tretyakov, 2003; Senior and Volakis, 1995; Kurushin and Nefedov, 1983; Collin, 2001; Weit, 1996) including plasma-like solid state (Ruibys and Tolutis, 1983) and space plasma (Weit, 1996) media. The plasma wave processes in the waveguide structures metal-semiconductor-dielectric, placed into the external magnetic field, were widely investigated (Ruibys and Tolutis, 1983; Maier, 2007; Tarkhanyan and Uzunoglu, 2006) in various frequency ranges, from radio to optical ones. Corresponding waves are applied in modern plasmonics and in non-destructive testing of semiconductor interfaces. It is of interest to realize the resonant interactions of volume and surface electromagnetic waves in these structures, so the simulations of the wave spectrum there are important. To describe such complex layered structures, it is very convenient and effective to use impedance approach (Tretyakov, 2003; Senior and Volakis, 1995; Kurushin and Nefedov, 1983). As a rule, impedance
boundary conditions are used, when the layer covering waveguide is thin (Senior and Volakis, 1995; Kurushin and Nefedov, 1983). One of the known exclusions is the impedance invariant imbedding model, the distinction to which of our new method has been mentioned above. Our new approach TIMEB, proposed in the present paper has the set of very attractive, for practical purposes, features. These features are: (i) the surface impedance characterizes cover layer of finite thickness, and this impedance is expressed analytically; (ii) the method allows an effective modelling of 3D beam propagating in the gyrotropic waveguiding structure; (iii) finally, if the considered waveguide can be modified by any external influence such as bias magnetic or electric fields, or by any extra wave or energy beams (such as acoustic or quasistatic fields etc.), the corresponding modification of the characteristics (phase and amplitude) of the electromagnetic VLF beam propagating in the waveguide structure can be modelled.

Our approach was targeting properly and is suitable for the farther important development which will allow to solve also the following problems, which continue the list presented above: (iv) the problem of the excitation of the waveguide by the waves incident on the considered structure from above could be solved as well with the slight modification of the presented model, with inclusion also ingoing waves; (v) consider a plasma-like system placed into the external magnetic field, such as the LAIM system (Grimalsky et al., 1999 a, b) or dielectric-magnetized semiconductor structure; then the waves radiated outside the waveguiding structure (such as helicons (Ruibys and Tolutis, 1983) or whistlers (Weit, 1996)) and the waveguide modes could be considered altogether; (vi) adequate boundary radiation conditions on the upper boundary of the covering layer are derived; and, based on this (and absence of ingoing waves), the leakage modes above the upper boundary of the structure (in other words, upper boundary of covering layer), will be searched with the farther development of the model, delivered in the present paper. Namely, the process of the leakage of the electromagnetic waves from the (opened) waveguide, then their transformation into magnetized plasma waves, propagating along magnetic field lines, and, possibly, excitation of the waveguiding modes by the waves incident on the system from external space (Walker, 1976), can be modelled as a whole. Such a modification can be measured, characterising the external fields and corresponding field sources, caused the above mentioned waveguide modification. Combining with the proper measurements of the phases and amplitudes of the electromagnetic waves, propagating in the waveguiding structures and leakage waves, the model possessing the above mentioned features can be used for searching, and even monitoring the external influences on the layered gyrotropic active artificial or natural media, for example microwave or optical waveguides or the system LAIM and WGEI, respectively.

The structure of the paper is as follows. In Sect. 2 formulation of the problem is presented. In Sect. 3 the algorithm is presented, including the determination of the conditions of radiation of the VLF waves/beams into the upper ionosphere/magnetosphere at the upper boundary, placed in the F region at the altitude (250-400) km; effective tensor impedance boundary conditions at the upper boundary (~ 85 km) of the effective WGEI; and finally the 3D model of the propagation of the VLF beam in the WGEI, which we call TIMEB, because in fact the beam method is combined with tensor impedance method. The questions on the mode presentation and leakage modes of VLF are discussed very briefly, because the corresponding details will be presented in the next papers. In Sect. 4, the results of numerical modelling are presented. In
Sect. 5, the discussion is presented, including an example of the qualitative comparison between the results of our theory and an experiment; and the future rocket experiment on the measurements of the characteristics of VLF signal, radiated from the VLF transmitter and propagating in the WGEI and penetrating into the upper ionosphere. Finally, conclusions are presented.

2 Formulation of the problem

The VLF electromagnetic (EM) waves with frequencies $f = 10 - 100 \text{ kHz}$ can propagate along the Earth’s surface for long distances >1000 km. The Earth’s surface of a high conductivity $z = 0$ and the ionosphere F-layer $z = 300 \text{ km}$ form the VLF waveguide, see Fig. 1. The propagation of the VLF electromagnetic radiation excited by a near-Earth antenna within the WGEI should be described by the full set of the Maxwell equations in the isotropic atmosphere $0 < z < 60 \text{ km}$, the approximately isotropic ionosphere D-layer $60 \text{ km} < z < 75 \text{ km}$, and the anisotropic E- and F- layers of the ionosphere, due to the geomagnetic field $H_\perp$, added by the boundary conditions at the Earth’s surface and at the F-layer.

Figure 1. The geometry of the anisotropic/gyrotropic waveguide. EM waves propagate in OX direction. $H_\perp$ is the external magnetic field. The (effective) WGEI for EM waves occupies the region $0 < z < L_{ISO}$, $L_{ISO} < L_z$. Anisotropic/gyrotropic media occupies the region $L_{ISO} < z < L_{max}$. Covering layer occupies the region $L_z < z < L_{max}$. WGEI includes isotropic region $0 < z < L_{ISO}$ and a part of anisotropic region $L_z < z < L_{max}$.

It is supposed that the anisotropic region is relatively small part of the WGEI, $(L_z - L_{ISO})/L_z \approx (0.1-0.2)$. At the upper boundary of covering layer ($z = L_{max}$) the radiation of EM to the external region ($z > L_{max}$) is accounted for with the proper boundary conditions. Integration of the equations describing the EM field propagation allows to obtain effective impedance boundary conditions at the upper boundary of effective WG ($z = L_z$). These boundary conditions effectively includes all the effect on the wave propagation of the covering layer and the radiation (at $z = L_{max}$) to the external region ($z > L_{max}$). $\theta$ is the angle between the directions of the vertical axis $\hat{z}$ and geomagnetic field $H_\perp$. The coordinate system $x''y''z''$ is connected with the geomagnetic field: $H_\perp$ is directed along $z''$ axis, lies in the plane $\Sigma$, and the planes $x''z''$ and $\Sigma$ coincide with each other.

3. Algorithm

We present here the algorithm of the new proposed method, staying in details only on the main subject of the present paper, in particular on the boundary conditions, impedance method and the method for the beam propagation in the WGEI. The
other parts of the method, connected with the mode presentation of the excitation of WGEI by a given current source and the reflection of the EM waves from the effective upper boundary of the WGEI and leakage of the EM waves from the WGEI to the upper ionosphere/magnetosphere, will be outlined here only very briefly and will become the subjects of the next papers.

3.1 Direct and inverse tensors characterizing the ionosphere

In the next subsections we will derive the formulas describing the transfer of the boundary conditions at the upper boundary \((z = L_{\text{max}})\), Fig. 1, resulting in the tensor impedance conditions at the upper boundary of the effective WGEI \((z = L_i)\). To make this, we need, firstly, to describe the tensors, characterizing the ionosphere. The monochromatic EM field is considered with the components of EM field \(\mathbf{E} \exp(i\omega t)\). The main goal is to transfer the EM boundary conditions from the upper ionosphere at the height \(L_z \sim 250 – 400\) km to the lower ionosphere \(L_z \sim 70 – 90\) km. The vertical axis is \(OZ\), the inclination angle of the geomagnetic field is \(\Theta\) (Fig. 1). The anisotropic medium is inhomogeneous along \(OZ\) axis only and is characterized by the tensor permittivity \(\varepsilon(\omega, z)\) or by the inverse tensor \(\hat{\beta}(\omega, z) = \varepsilon^{-1}(\omega, z); \hat{E} = \hat{\beta}(\omega, z) \cdot \hat{D}\), where \(\hat{D}\) is the electric induction. Below the absolute units are utilized. The expressions for the components of the effective permittivity of the ionosphere are in the coordinate frame \(X'YZ'\) where \(OZ'\) axis is aligned along the geomagnetic field \(\vec{H}_e\):

\[
\varepsilon^* = \begin{pmatrix}
\varepsilon_x & \varepsilon_y & 0 \\
-\varepsilon_y & \varepsilon_x & 0 \\
0 & 0 & \varepsilon_z
\end{pmatrix}, \quad \varepsilon^*_x = \frac{1}{\omega_\psi^2} \left(\frac{\omega_\psi^2 - \omega - i\nu_p}{\omega_\psi^2 + \omega - i\nu_p}\right) - \frac{\omega_\psi^2 - \omega - i\nu_i}{\omega_\psi^2 + \omega - i\nu_i} + \frac{\omega_\psi^2 - \omega - i\nu_e}{\omega_\psi^2 + \omega - i\nu_e}
\]

\[
\varepsilon^*_y = \frac{1}{\omega_\psi^2} \left(\frac{\omega_\psi^2 - \omega - i\nu_p}{\omega_\psi^2 + \omega - i\nu_p}\right) - \frac{\omega_\psi^2 - \omega - i\nu_i}{\omega_\psi^2 + \omega - i\nu_i} + \frac{\omega_\psi^2 - \omega - i\nu_e}{\omega_\psi^2 + \omega - i\nu_e}
\]

\[
\varepsilon^*_z = \frac{1}{\omega_\psi^2} \left(\frac{\omega_\psi^2 - \omega - i\nu_p}{\omega_\psi^2 + \omega - i\nu_p}\right) - \frac{\omega_\psi^2 - \omega - i\nu_i}{\omega_\psi^2 + \omega - i\nu_i} + \frac{\omega_\psi^2 - \omega - i\nu_e}{\omega_\psi^2 + \omega - i\nu_e}
\]

(1)

Here \(\omega_\psi, \omega_p, \omega_i, \omega_e\) are plasma and cyclotron frequencies for electrons and ions respectively; \(m_e, m_i, v_e, v_i\) are the masses and the collision frequencies. The expressions of the components of \(\varepsilon(\omega, z)\) are obtained from (1) by means of multiplication with the rotation matrices (Spiegel, 1959). In the case of a medium with a scalar conductivity \(\sigma\), like the lower ionosphere or atmosphere, the effective permittivity (1) reduces to the scalar: \(\varepsilon = 1 - 4\pi\sigma/\omega\)

3.2 The equations for the EM field and upper boundary conditions

The EM field depends on the horizontal coordinate \(x\) as \(-\exp(-ik_x x)\). Generally, \(k_x \leq k_0\), where \(k_0 = \omega/c\). In simulations of VLF beam propagation, we put \(k_x = k_0\). In the case of searching VLF waveguide modes \(k_x\) is slightly complex and should be calculated from boundary conditions at the Earth’s surface and upper surface of the effective WGEI.

The Maxwell equations are:
In eq. (2), $E_z = \beta_1 D_x + \beta_2 D_y + \beta_3 D_z$ etc. All the components of the EM field can be represented through the horizontal components of the magnetic field $H_x$, $H_y$, and the following equations for these components have been derived:

\[
\frac{\partial H_x}{\partial z} = \frac{\partial H_y}{\partial z}, \quad \frac{\partial H_z}{\partial z} = \frac{\partial H_x}{\partial z}, \quad -\frac{\partial H_y}{\partial z} = \frac{\partial H_z}{\partial z} \quad (2)
\]

\[
-\frac{\partial E_x}{\partial z} = -ik_0 H_y, \quad \frac{\partial E_y}{\partial z} = -ik_0 H_x, \quad -\frac{\partial E_z}{\partial z} = -ik_0 H_z
\]

The expressions for the horizontal components of the electric field $E_x$, $E_y$ are:

\[
E_x = \frac{i}{k_0} \left( \frac{\partial H_x}{\partial z} - \frac{\partial H_y}{\partial z} \right) H_z + \frac{\beta_{12} + \beta_{13} \beta_{23} \frac{k_2^2}{k_0^2} \frac{k_3^2}{k_0^2} - \beta_{23} \frac{k_2^4}{k_0^4} - \beta_{13} \beta_{23} \frac{k_3^4}{k_0^4}}{1 - \beta_{23} \frac{k_2^2}{k_0^2} - \beta_{13} \beta_{23} \frac{k_3^2}{k_0^2}} \right)
\]

\[
E_y = \frac{i}{k_0} \left( \frac{\partial H_y}{\partial z} - \frac{\partial H_z}{\partial z} \right) H_x + \frac{\beta_{12} + \beta_{13} \beta_{23} \frac{k_2^2}{k_0^2} \frac{k_3^2}{k_0^2} - \beta_{23} \frac{k_2^4}{k_0^4} - \beta_{13} \beta_{23} \frac{k_3^4}{k_0^4}}{1 - \beta_{23} \frac{k_2^2}{k_0^2} - \beta_{13} \beta_{23} \frac{k_3^2}{k_0^2}} \right)
\]

In the region $z \geq L_{max}$, the upper ionosphere is assumed weakly inhomogeneous, and the geometric optics approximation is valid in the VLF range there. Note that such an approximation is invalid at the upper boundary of the effective VLF WGEI at $80 - 90$ km because of the great inhomogeneity of the ionosphere in the vertical direction within $E$-layer. These circumstances determine the choice of the upper boundary $z = L_{max}$ ~ (250-400) km, where the conditions of the radiation are
formulated. The dispersion equation connected the wave numbers and the frequency of the outgoing waves has been got from eqs. (3), where \( H_{x,y} \sim e^{-i k_z z} \), while the derivatives like \( \partial b_1 / \partial z \) and the inhomogeneity of the media are neglected:

\[
(k_2^2 - k^2)(1 - \beta_{22} k_z^2 / k_0^2) \left( (\beta_{11} - \beta_{22} k_z^2 / k_0^2 + \beta_{12} \cdot \beta_{21}) k_z^2 + ((\beta_{11} + \beta_{22})(1 - \beta_{22} k_z^2 / k_0^2) + \right.
\]
\[
+ \beta_{12} \cdot \beta_{21} k_z^2)((1 - \beta_{22} k_z^2 / k_0^2)(1 - \beta_{22} k_z^2 / k_0^2) - k_4^2) - k_z^2 \beta_{33} \cdot \beta_{21}) \right) -
\]
\[
- \left( \beta_{11} k_z^2 + \beta_{22} k_z^2 \right) \left( \beta_{11} k_z^2 \beta_{22} k_z^2 \right) = 0
\]

Thus, generally Eq. (5) which determines the wave numbers for the outgoing waves is of the 4th order (Weit 1996). The boundary conditions at the upper boundary \( z = L_{max} \) within the ionosphere \( F^- \) layer are the absence of the ingoing waves, i.e. the outgoing (radiated) waves are present only. Two roots should be selected that possess the negative imaginary parts \( \text{Im}(k_z) < 0 \), i.e. the outgoing waves dissipate upwards. But in the case of VLF waves some simplification can be used. Namely, the expressions for the wave numbers \( k_{12} \) are obtained from eqs. (3), where the dependence on \( x \) is neglected: \( |k_{12}| \gg k_0 \). This approximation is valid within \( F^- \) layer where the first outgoing wave corresponds to the whistler of small dissipation, the second one to the highly dissipating slow wave. The EM field components, which are necessary to formulate the boundary conditions for eqs. (3a, b) at \( z \geq L_{max} \), can be presented as:

\[
H_z = A_1 e^{-i k_z z} + A_2 e^{-i k_z z}, \quad H_y = \alpha_1 A_1 e^{-i k_z z} + A_2 e^{-i k_z z}
\]

In the relations (6), \( \tilde{z} = z - L_z \), eqs. (3) are simplified there in the approximation described above:

\[
\beta_{22} \frac{\partial^2 H_z}{\partial z^2} - \beta_{22} \frac{\partial^2 H_y}{\partial z^2} + k_z^2 H_z = 0, \quad \beta_{11} \frac{\partial^2 H_y}{\partial z^2} \quad \beta_{12} \frac{\partial^2 H_z}{\partial z^2} + k_z^2 H_y = 0
\]

Again, the solution of eqs. (7) is searched as: \( H_{x,y} \sim e^{-i k_z z} \). The following equation has been obtained to get the wave numbers \( k_{12} \) from eqs. (7):

\[
(\beta_{22} + \beta_{11}) k_z^2 + \beta_{11} \beta_{22} - \beta_{12} \beta_{21} = 0, \quad k_z^2 = \frac{k_2^2}{k_1^2}
\]

Therefore as it follows from eq. (8),

\[
k_{12}^2 = \frac{\beta_{11} + \beta_{22}}{2} \left( \frac{\beta_{11} + \beta_{22}}{2} + \beta_{12} \beta_{21} \right) \right)^{1/2}, \quad \alpha_1 = \beta_{22} - k_z^2 \beta_{22} = \frac{\beta_{12} - \beta_{22}}{\beta_{21}}, \quad \alpha_2 = \frac{\beta_{12} - \beta_{22}}{\beta_{21}} k_{12}^2 = \frac{k_2^2}{k_1^2}
\]

The signs of \( k_{12} \) have been chosen from the condition \( \text{Im}(k_{12}) < 0 \). From Eqs. (5) at the upper boundary \( z = L_{max} \) the following relations are valid:

\[
H_z = A_1 + \alpha_1 A_2, \quad H_y = \alpha_1 A_1 + A_2
\]

As it follows from eq. (10),

\[
A_1 = \Delta^{-1}(H_y - \alpha_1 H_z), A_2 = \Delta^{-1}(H_y - \alpha_1 H_z); \Delta = 1 - \alpha_1 \alpha_2
\]
Thus, it is possible to exclude the amplitudes of the outgoing waves $A_{1,2}$ from Eqs. (9). As a result, at $z = L_{\text{max}}$, the boundary conditions are rewritten in terms of $H_x, H_y$ only:

\[
\frac{\partial H_x}{\partial z} = -i(k_{i1}A_1 + k_{i2}A_2) = -\frac{i}{\Delta} \left( (k_{i1} - \alpha \alpha A_{i2}H_y + \alpha \alpha (k_{i2} - k_z))H_y \right) \\
\frac{\partial H_y}{\partial z} = -i(k_{i1}A_1 + k_{i2}A_2) = -\frac{i}{\Delta} \left( (k_{i2} - \alpha \alpha A_{i2}H_y + \alpha \alpha (k_{i1} - k_z))H_x \right)
\] (12)

The relations (12) are the upper boundary conditions of the radiation for the boundary $z=L_{\text{max}}=\sim(250-400)$ km. Then these conditions will be transformed/recalculated using the analytical-numerical recurrent procedure into equivalent impedance boundary conditions at $z=L_{\text{z}}=\sim(70-90)$ km.

Note that in the “whistler/VLF approximation”, valid at frequencies $\sim 10$ kHz, one can get for the F region of the ionosphere. In this approximation and accounting for that $k_x = 0$, we find, using eqs. (5), (8), (9) that dispersion equation takes the form

\[ k_x^2 k_y^2 = k_o^2 \] (13)

where $k^2 = k_x^2 + k_y^2 = k_{i1}^2 + k_{i2}^2$; $k_{i1}$ and $k_{i2}$ are the components of wave number, transverse and longitudinal respectively to geomagnetic field, respectively. For the F region of the ionosphere, where $v_e \ll \omega \ll \omega_{pi}$, eq. (13) reduces to the standard form of whistler dispersion equation $|k_x| k_y = k_o \cos \theta$; $g = -\omega^2 / (\omega \omega_{pi})$; $\omega = c^2 k |k_o| \omega_{pi}$; in a special case of the waves, propagating exactly along geomagnetic field, $k_x = 0$, one obtain, for the propagating whistler waves, well-known dispersion dependence (Artcimovich and Sagdeev, 1979) $\omega = c^2 k_o^2 (\omega_{pi} / \omega_o^2)$. Coming back to our problem and accounting for that in our case we can reasonably put $k_x = 0$, eq. (13) reduces to $k_{i1}^2 \cos^2 \theta = k_{o1} g^2$. As a result, we get $k_{i1} = \sqrt{g / \cos \theta k_o}$, $k_{i2} = -\sqrt{g / \cos \theta k_o}$, and then, similarly to the relations (12), the boundary conditions can be presented, in terms of the tangential components of electric field, particularly in the form:

\[
\frac{\partial E_x}{\partial z} + \frac{\partial E_y}{\partial z} = 0; \quad \vec{U} = \begin{bmatrix} E_x \cr E_y \cr \end{bmatrix}; \quad \vec{B} = \frac{1}{2} \frac{g}{\cos \theta k_o} \begin{bmatrix} 1+i & 1-i \\
1+i & 1+i \end{bmatrix}
\] (14)

Conditions (12) or (14) are the conditions of radiation (absence of ingoing waves) formulated at the upper boundary $z=L_{\text{max}}$ and suitable for the determination of the energy of the wave leaking from the WGEI into the upper ionosphere/magnetosphere. Let us emphasize again that the formulas expressing the boundary conditions of the radiation (more accurately speaking, an absence of incoming waves, what is the consequence to the causality principle) (12), (14) are obtained as a result of limiting pass by the small parameter $k_x/k_o \to 0$ in eq. (5). Note that in spite of disappearance of the dependence of these boundary conditions explicitly on $k_o$, the dependence of the characteristics of the wave propagation process on $k_o$, as a whole, is accounted for, and all results are still valid for the description of the wave beam propagation in the WGEI along the horizontal axis $x$ with finite $k_x - k_{i1}$. 
### 3.3 Equivalent Tensor Impedance Boundary Conditions at the Upper Boundary \( z=L_z \) of the Effective WGEI

The tensor impedance at the upper boundary of the effective WGEI \( z=L_z \), Fig. 1, is obtained by means of recalculating to the level \( z=L_z \approx 80 \) – \( 90 \) km of the conditions of radiation (12) or (14), formulated at the upper boundary, placed in the F region of the ionosphere, at \( z=L_{\text{max}} \approx (250-400) \) km.

The main idea of the effective tensor impedance method is the unification of the analytical and numerical approaches and the derivation of the proper impedance boundary conditions without any approximation of the “thin cover layer”, used in the majority of an effective impedance approaches previously, applied either for artificial or natural layered gyrotrropic structures, see, f.e. (Tretyakov, 2003; Senior and Volakis, 1995; Kurushin and Nefedov, 1983; Alperovich and Fedorov, 2007). There is one known exception, namely invariant imbedding impedance method (Shalashov and Gospodchikov, 2010; Kim and Kim, 2016). The comparison of our method with the invariant imbedding impedance method will be presented in the end of this subsection. Eqs. (3), jointly with the boundary conditions (12), have been solved by finite differences. Outline here the main ideas and the steps of the derivations of the corresponding formulas.

The derivatives in Eqs. (3) are approximated as

\[
\frac{\partial}{\partial z} \left( C(z) \frac{\partial H_j}{\partial z} \right) \approx \frac{1}{h} \left( C(z_{j+1}) \left( \frac{H_j}{h} \right)_{j+1} - C(z_{j-1}) \left( \frac{H_j}{h} \right)_{j-1} \right).
\]

In eq. (15), \( z_{j+0.5} = h \cdot (j + 0.5) \). In eqs. (10) the approximation is \( \partial H_j / \partial z = [(H_j)_z - (H_z)_{j-1}] / h \). Here \( h \) is the discretization step along \( OZ \) axis, \( N \) is the total number of the nodes. At each step \( j \) the difference approximations of Eqs. (3) take the form:

\[
\hat{a}^{(1)}_j \hat{H}_{j-1} + \hat{a}^{(8)}_j \hat{H}_j + \hat{a}^{(4)}_j \hat{H}_{j+1} = 0
\]

where \( \hat{H}_j = \begin{pmatrix} H_j \\ H_z \end{pmatrix} \). \( j = N-1, N-2, \ldots, 1 \), \( z_j = h \cdot j \), \( L_z = h \cdot N \). The expressions for the matrix coefficients in Eq. (16) are complicated; they are given in Appendix. The set of the matrix eqs. (16) has been solved by the method called factorization, or elimination, or matrix sweep (method) [Samarskii, 2001]. Namely, it is possible to write down:

\[
\hat{H}_j = \hat{b}_j \cdot \hat{H}_{j-1}, \quad j = N, \ldots, 1
\]

\[
H_{a1} = b_{11} \cdot H_1 + b_{21} \cdot H_z; \quad H_{a2} = b_{12} \cdot H_1 + b_{22} \cdot H_z; \quad H_1 = H_{a1}; \quad H_z = H_{a2}
\]

This method is in fact a variant of the Gauss elimination method for the matrix 3-diagonal set of the Eqs. (16). The value of \( \hat{b}_j \) has been obtained from the boundary conditions (12). Namely, they can be rewritten as:
\[ \hat{a}_n^{(+)} \cdot \hat{H}_{n-1} + \hat{a}_n^{(0)} \cdot \hat{H}_n = 0 \]  
(18)

Therefore \( \hat{b}_j = -(\hat{a}_n^{(0)})^{-1} \cdot \hat{a}_n^{(+)} \). Then the matrices \( \hat{b}_j \) have been computed sequentially down till the desired value of \( z = L_\ell = h \cdot N_\ell \), where the impedance boundary conditions are assumed to be applied. At each step the formulas for \( \hat{b}_j \) follow from (16), (17) and take the form

\[ (\hat{a}_j^{(0)} + \hat{a}_j^{(+)} \cdot \hat{b}_{j+1} \cdot \hat{H}_j = -\hat{a}_j^{(+)} \cdot \hat{H}_{j+1} \]  
(19)

Therefore accounting for (17), we obtain \( \hat{b}_j = -(\hat{a}_j^{(0)} + \hat{a}_j^{(+)} \cdot \hat{b}_{j+1})^{-1} \cdot \hat{a}_j^{(+)} \). The derivatives in eqs. (4) have been approximated as:

\[ \frac{\partial H_j}{\partial z} = \frac{(H_{j+1} - H_j)}{h} \cdot (b_{j+1} - 1) \cdot (H_{j+1} - H_j) + b_{j+1} \cdot (H_{j+1} - H_j) \]  
(20)

Note that a result of this discretization, only the values at the grid level \( N_\ell \) are included into the numerical approximation of the derivatives \( \partial H_{j,\ell} / \partial z \) at \( z = L_\ell \). We determine tensor impedance \( Z \) in particular at \( z = L_\ell \sim 85 \text{ km} \) such as the tensor value containing in the following relations, all of which are related to the corresponding altitude (in other words, to the grid with number \( N_\ell \), corresponding to this altitude):

\[ \hat{n} \times \hat{E} = \hat{Z} \cdot \hat{H}_j, \hat{n} = (0,0,1); \text{ or } E_x = Z_{j,0} H_x + Z_{j,1} H_y; E_y = -Z_{j,0} H_x - Z_{j,1} H_y \]  
(21)

The equivalent tensor impedance is obtained, in fact, using two-step procedure. (1) We obtain the matrix \( \hat{b}_j \) using the set of equations (3a, b) with the boundary conditions (12) and the procedure (17)-(19) described above. (2) Put the expressions (21) with tensor impedance into the left parts and the derivatives \( \partial H_{j,\ell} / \partial z \) in the form (20) into the right parts of eqs. (4).

Equating in the left and right parts of the two obtained equations coefficients at \( H_x \), \( H_y \) respectively, we obtain the analytical expressions for the components of the tensor impedance at \( z = L_\ell \):

\[ Z_{11} = -i \frac{k_0 \hbar}{k_0} \left( \frac{\beta_{21}}{1 - \beta_{22} k_0^2} \cdot b_{21} - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot (b_{11} - 1) \right), \]

\[ Z_{12} = -i \frac{k_0 \hbar}{k_0} \left( \frac{\beta_{21}}{1 - \beta_{22} k_0^2} \cdot \frac{\partial H_{11}}{\partial z} \cdot (b_{22} - 1) - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot b_{21} - k_0 \hbar \cdot \frac{\beta_{23}}{1 - \beta_{22} k_0^2} \right), \]

\[ Z_{22} = \frac{i}{k_0 \hbar} \left( \frac{k_0^2 \beta_{21}}{1 - \beta_{22} k_0^2} \cdot b_{22} - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot (b_{11} - 1) \right), \]

\[ Z_{23} = -i \frac{k_0 \hbar}{k_0} \left( \frac{\beta_{21}}{1 - \beta_{22} k_0^2} \cdot \frac{\partial H_{11}}{\partial z} \cdot (b_{22} - 1) - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot b_{21} - k_0 \hbar \cdot \frac{\beta_{23}}{1 - \beta_{22} k_0^2} \right), \]

\[ Z_{31} = \frac{i}{k_0 \hbar} \left( \frac{k_0^2 \beta_{21}}{1 - \beta_{22} k_0^2} \cdot b_{22} - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot (b_{11} - 1) \right), \]

\[ Z_{32} = -i \frac{k_0 \hbar}{k_0} \left( \frac{\beta_{21}}{1 - \beta_{22} k_0^2} \cdot \frac{\partial H_{11}}{\partial z} \cdot (b_{22} - 1) - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot b_{21} - k_0 \hbar \cdot \frac{\beta_{23}}{1 - \beta_{22} k_0^2} \right), \]

\[ Z_{33} = \frac{i}{k_0 \hbar} \left( \frac{k_0^2 \beta_{21}}{1 - \beta_{22} k_0^2} \cdot b_{22} - \frac{\beta_{22}}{1 - \beta_{22} k_0^2} \cdot (b_{11} - 1) \right), \]
The proposed method of the transfer of the boundary conditions from the ionosphere $F$-layer $L_{\text{max}} = 250 \text{ – } 400 \text{ km}$ into the lower part of the $E$-layer $L_z = 80 \text{ – } 90 \text{ km}$ is stable and easily realizable, when compared with some alternative approaches based on the invariant imbedding methods (Shalashov and Gospodchikov, 2010; Kim and Kim, 2016). The stability of our method is due to the stability of the Gauss elimination method when the coefficients at the central diagonal are dominating; the last is valid for the ionosphere with electromagnetic losses where the absolute values of the permittivity tensor are big. The application of the proposed method of the matrix sweep in the media without losses may require the utilization of the Gauss method with the choice of the maximum element, to ensure the stability. But, as our simulations (not presented here) demonstrated, for the electromagnetic problems in the frequency domain the simple Gauss elimination and one with the choice of the maximal element gives the same results. The accumulation of errors may occur in evolutionary problems in the time domain, when the Gauss method should be applied sequentially many times. The use of the independent functions $H_x$, $H_y$ in Eqs. (3) seems natural, as well as the transfer (17a), because the impedance conditions are the expressions of the electric $E_x, E_y$ through these magnetic components $H_x, H_y$ at the upper boundary of the VLF waveguide $80 \text{ – } 90 \text{ km}$. The naturally chosen direction of the recalculation of the upper boundary conditions from $z=L_{\text{max}}$ to $z=L_z$, i.e from upper layer with large impedance value to lower altitude layer with relatively small impedance value, provides, at the same time, the stability of the simulation procedure. The obtained components of the tensor impedance are small, $|Z_{ab}| \leq 0.1$. This determines the choice of the upper boundary $z=L_z$ of the effective WGEI. Due to small enough impedance, EM waves incident from below on this boundary reflect effectively back. Therefore, the region $0 \leq z \leq L_z$ indeed can be presented as an effective WGEI. Then such a waveguide includes not only lower up to $L_{\text{ISO}} \sim (65\text{-}75) \text{ km}$ with rather small losses, but also thin dissipative and anisotropic/gyrotropic layer between 75 and 85-90 km.

Finally, the main differences and at the same time advantages of the proposed tensor impedance method from the known method of the impedance recalculating, in particular invariant imbedding methods (Shalashov and Gospodchikov, 2010; Kim and Kim, 2016) are the following. (i) In distinction to invariant imbedding method, our method is a direct method of the recalculation of tensor impedance, and the corresponding tensor impedance is determined analytically, see eqs. (22). (ii) Our method, for the media without non-locality, does not need a solution of integral equation(s), as in invariant imbedding method. (iii) The proposed tensor impedance method does not need the revealing of forward and reflected waves. Moreover, even the conditions of the radiation (12) at the upper boundary $z=L_{\text{max}}$ are determined through the total field components $H_{ex}$, what makes the proposed procedure technically much less cumbersome and practically much more convenient. (iv) At the same time, the procedure is very effective and computationally stable, as it is explained above in this subsection. As it is already mentioned, for the very low-loss systems, the required level of stability can be achieved with the modification based on the choice of the maximal element for matrix inversion.
3.4 Propagation of the Electromagnetic Waves in the Gyrotropic Waveguide and the TIMEB Method

Use, as the independent functions, the transverse components $E_y$, $H_y$. The goal is to derive the equations for the slowly varying amplitudes $A(x,y,z)$, $B(x,y,z)$ of the VLF beams included corresponding field components:

$$ E_y = \frac{1}{2} A(x,y,z) \cdot e^{i\omega - ik_y z} + c.c., \quad H_y = \frac{1}{2} B(x,y,z) \cdot e^{i\omega - ik_y z} + c.c. \quad (23) $$

Note that in this case it should be $k_x = k_0$, because the beam propagates in the WGEI, the main part of which is occupied by the atmosphere and lower ionosphere (D region), rather closed to free space by its electromagnetic parameters. The presence of a thin anisotropic and dissipative layer belonging to the E region of the ionosphere causes, altogether with the impedance boundary condition the proper $z$ dependence of $B(x,y,z)$. Using (21), (22), it is possible to write down the boundary conditions at the height $z = L_z$ for the slowly varying amplitudes $A(x,y,z)$, $B(x,y,z)$ of the transverse components $E_y$, $H_y$:

$$ A - \frac{i}{k_0} Z_{11} \frac{\partial A}{\partial z} + Z_{12} \cdot B \equiv 0, \quad \gamma_{12} \cdot A + \frac{i}{k_0} Z_{21} \frac{\partial A}{\partial z} + (\hat{P}_{13} - Z_{22}) \cdot B + \frac{i}{k_0} \hat{P}_{13} \frac{\partial B}{\partial z} \equiv 0 \quad (25) $$

Let us derive the evolution equations for the slowly varying amplitudes $A(x,y,z)$, $B(x,y,z)$ of the VLF beams. Below the monochromatic beams are considered, so the frequency $\omega$ is fixed and the amplitudes do not depend on time $t$. Search the solutions for the EM field as $\vec{E}, \vec{H} \sim \exp(\mathbf{i} \omega t - \mathbf{i} k_x x - \mathbf{i} k_y y)$. The Maxwell equations are written down as

$$ -\mathbf{i} k_y H_x - \frac{\partial H_x}{\partial z} = \mathbf{i} k_0 D_x, \quad \frac{\partial H_x}{\partial z} + \mathbf{i} k_y H_x = \mathbf{i} k_0 D_x, \quad -\mathbf{i} k_y H_y - \mathbf{i} k_0 H_y = \mathbf{i} k_0 D_y $$

$$ -\mathbf{i} k_x E_y - \frac{\partial E_y}{\partial z} = -\mathbf{i} k_0 H_y, \quad \frac{\partial E_y}{\partial z} + \mathbf{i} k_x E_y = -\mathbf{i} k_0 D_y, \quad -\mathbf{i} k_x E_x + \mathbf{i} k_0 E_x = -\mathbf{i} k_0 H_x \quad (26) $$

Here $D_x = e_{11} E_x + e_{12} E_y + e_{13} H_y$. From Eqs. (21) it is possible to get the expressions for $E_y$, $E_x$ through $E_y$, $H_y$:

$$ E_y = \frac{1}{\Delta} \left[ e_{11} (e_{11} - e_{12} - e_{13}) \cdot (e_{11} - \frac{k_x}{k_0} H_x) E_x - \frac{k_y}{k_0} \frac{\partial H_x}{\partial z} + \frac{k_y}{k_0} \frac{\partial H_x}{\partial z} + \mathbf{i} k_0 e_{11} E_y + \mathbf{i} k_0 e_{13} E_x \right] $$

$$ E_x = \frac{1}{\Delta} \left[ e_{12} (e_{12} + \frac{k_x}{k_0}) - e_{13} (e_{13} - \frac{k_x}{k_0}) \cdot (e_{12} - \frac{k_x}{k_0}) E_x - \frac{k_y}{k_0} \frac{\partial H_x}{\partial z} - \mathbf{i} k_0 \frac{k_y}{k_0} (e_{11} - e_{13}) H_x - \frac{k_y}{k_0} \cdot (e_{12} - \frac{k_x}{k_0}) \frac{\partial E_y}{\partial z} \right] \quad (27) $$

In eq. (27), $\Delta = (e_{11} - \frac{k_x^2}{k_0^2}) \cdot (e_{13} - \frac{k_x^2}{k_0^2}) - e_{12} \cdot e_{13}$. The equations for $E_y$, $H_y$ obtained from the Maxwell equations are:
After substitution of expressions (27) for $E_x$, $E_z$ into Eqs. (28) the coupled equations for $E_y$, $H_y$ only can be got. Namely, the expansion should be used: $k_x = k_{0y} + \delta k_x$, $|\delta k_x| << k_{0y}$, also $|k_y| << k_{0y}$. Then the correspondence should be applied (Weiland and Wilhelmsson 1977):

$$\begin{align*}
-i \cdot \delta k_x & \rightarrow \frac{\partial}{\partial x}, \quad -i \cdot k_y \rightarrow \frac{\partial}{\partial y}.
\end{align*}$$

The expansions should be till the quadratic terms with respect to $k_y$ and the linear terms with respect to $\delta k_x$. As a result, the parabolic equations (Levy 2000) for the slowly varying amplitudes $A$ and $B$ have been derived. In the atmosphere and the lower ionosphere, where the effective permittivity reduces to a scalar $\varepsilon(\alpha z)$, they are independent:

$$\begin{align*}
\frac{\partial A}{\partial x} + \frac{i}{2k_0} \left( \frac{\partial^2 A}{\partial y^2} - \frac{\partial^2 A}{\partial z^2} \right) + \frac{i k_b}{2} \varepsilon_0 (\varepsilon - 1) A &= 0, \\
\frac{\partial B}{\partial x} + \frac{i}{2k_0} \left( \frac{\partial^2 B}{\partial y^2} - \frac{\partial^2 B}{\partial z^2} \right) + \frac{i k_b}{2} \varepsilon_0 (\varepsilon - 1) B &= 0.
\end{align*}$$

Here $\beta = \varepsilon^{-1}$. Accounting for the presence of gyrotropic layer near the and the presence of tensor impedance boundary conditions at the upper boundary $z = L_z$ of the VLF waveguide, the equations for the slowly varying amplitudes in general case are coupled and possess a complicated form:

$$\begin{align*}
\frac{\partial A}{\partial x} + \frac{i}{2k_0} \left( \frac{\partial^2 A}{\partial y^2} + \frac{\partial^2 A}{\partial z^2} \right) + \frac{ik_0}{2} (\varepsilon_0 - 1) A + \gamma_{11} \frac{\partial B}{\partial z} + \gamma_{12} B &= 0, \\
\frac{\partial B}{\partial x} + \frac{i}{2k_0} \left( \frac{\partial^2 B}{\partial y^2} + \frac{\partial^2 B}{\partial z^2} \right) + \frac{ik_0}{2} (\beta_0 - 1) B &= 0.
\end{align*}$$

In Eq. (30b),

$$\begin{align*}
\gamma_{11} &= \frac{\varepsilon_0 - \varepsilon_{21} - \varepsilon_{31} - \varepsilon_{12}}{\Delta}, \quad \gamma_{12} = \frac{\varepsilon_0 - \varepsilon_{21} - \varepsilon_{31} - \varepsilon_{12}}{\Delta}, \\
\gamma_{21} &= \frac{\varepsilon_0 - \varepsilon_{21} - \varepsilon_{31} - \varepsilon_{12}}{\Delta}, \quad \gamma_{22} = \frac{\varepsilon_0 - \varepsilon_{21} - \varepsilon_{31} - \varepsilon_{12}}{\Delta}, \\
\beta_{11} &= \frac{\varepsilon_0 - \varepsilon_{21} - \varepsilon_{31} - \varepsilon_{12}}{\Delta}, \quad \beta_{12} = \frac{\varepsilon_0 - \varepsilon_{21} - \varepsilon_{31} - \varepsilon_{12}}{\Delta}.
\end{align*}$$

Eqs. (30b) reduce to Eqs. (30a) when the effective permittivity is scalar. At the Earth’s surface $z = 0$ the impedance conditions reduce, due to a finite conductivity of the Earth, to the form:

$$E_y = Z_e H_y, \quad E_z = -Z_e H_y, \quad Z_e = \left( \frac{\varepsilon_0}{4\pi\sigma_e} \right)^{1/2}.$$
Here $\sigma_e \sim 10^8 \text{ s}^{-1}$ is the Earth’s conductivity. The boundary conditions (31a) at the Earth’s surface, where $Z_{in} = Z_{in} = Z_e$, $Z_{in} = Z_{in} = 0$, $\beta_{in} = \varepsilon(z = 0)^{2/3}$, $\gamma_{in} = 0$, $\beta_{in} = 0$, can be rewritten as

$$E_y + \frac{i}{k_0} Z_e \frac{\partial E_x}{\partial z} = 0, \quad \frac{i}{\varepsilon(z = 0) k_0} \frac{\partial H_y}{\partial z} + \frac{i}{k_0} Z_e H_y = 0$$

(31b)

The Eqs. (30) added by the boundary conditions (25) at the upper boundary of the VLF waveguide $z = L_z$, and by the boundary conditions at the Earth’s surface (31b) are used below to simulate the VLF wave propagation. The surface impedance of the Earth has been calculated from the Earth’s conductivity. The initial conditions to the problem (30), (25), (31b) are

$$A(x = 0, y, z) = 0, \quad B(x = 0, y, z) = B_0 \exp\left(-\left[(y - 0.5 L_y)/y_0\right]^{2n}\right) \exp\left(-\left[(z - z_1)/z_0\right]^{2n}\right), \quad n = 2$$

(32)

The size of the computing region along $OY$ axis is $L_y \sim 1000 \text{ km}$. Because the gyrotropic layer is relatively thin and is placed at the upper part of the VLF waveguide, whereas the beams are excited near the Earth’s surface, the wave diffraction in this gyrotropic layer along $OY$ axis is quite small, i.e. the terms with $\partial^2 A/\partial y^2$, $\partial^2 B/\partial y^2$ are small there. Contrary to this, the wave diffraction is very important in the atmosphere in the lower part of the VLF waveguide near the Earth’s surface. To solve the problem of the beam propagation, the method of splitting with respect to physical factors has been applied (Samarskii 2001).

Namely, the problem has been approximated by the finite differences:

$$\tilde{C} = \begin{pmatrix} A \\ B \end{pmatrix}, \quad \frac{\partial \tilde{C}}{\partial x} + \hat{L}_y \tilde{C} + \hat{L}_z \tilde{C} = 0$$

(33)

In the terms $\hat{L}_y, \hat{C}$ the derivatives with respect to $y$ are included, whereas all another terms are included into $\hat{L}_z \tilde{C}$. Then the following fractional steps have been applied, the 1st one is along $y$, the 2nd one is along $z$:

$$\frac{\tilde{C}^{p+1/2} - \tilde{C}^p}{h_y} + \hat{L}_y \tilde{C}^{p+1/2} = 0, \quad \frac{\tilde{C}^{p+1} - \tilde{C}^{p+1/2}}{h_z} + \hat{L}_z \tilde{C}^{p+1/2} = 0$$

(34)

The region of simulation is $0 < x < L_x = 1000 – 2000 \text{ km}$, $0 < y < L_y = 2000 – 3000 \text{ km}$, $0 < z < L_z = 80 – 90 \text{ km}$. The numerical scheme (34) is absolutely stable. Here $h_y$ is the step along $OX$ axis, $x_p = p h_x$, $p = 0, 1, 2, \ldots$. This step has been chosen from the condition of the independence of the simulation results on diminishing $h_y$. 

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Figure 2. The rotation of the local Cartesian coordinate frame at each step along the Earth’s surface $h_x$ on a small angle $\delta \varphi \approx \Delta x/R_E$, radians, while $\Delta x = h_x$. The following strong inequalities are valid $h_x \ll L_z \ll R_E$. The Earth’s surface is at $Z = 0$.

Under the simulations at each step along $OX$ axis the correction due to the Earth’s curvature has been inserted in adiabatic manner, namely the rotation of the local coordinate frame $XOZ$. Because the step along $x$ is small $h_x \sim 1 \text{ km} \ll L_o$, this correction of the \( \mathcal{C} \) results in the multiplier $\exp(-ik_0 \cdot \delta x)$, where $\delta x = z(h_x/R_E)$, $R_E >> L_o$ is the Earth’s radius, see Fig. 2 and the capture to this figure. At the distances $x \leq 1000 \text{ km}$ the results of simulations do not depend on the insertion of this correction, whereas at higher distances some quantitative differences occur. Namely, the VLF beam propagates more closely to the upper boundary of the waveguide.

3.5 The Modes of the VLF Waveguide. Reflection from the Upper Effective Boundary of VLF Waveguide and Leakage Modes: Idea of Synthetic Full Wave Electromagnetic-Complex Geometrical Optics (FWEM-CGO) Approach for Magnetized Plasma

Our model, in general, needs the consideration of the excitations of the waveguide modes by means of current sources such as dipole-like VLF radio source and lightning discharges. Then, we will present the results of the reflection of the waves incident on the upper boundary ($z=L_z$) of the effective WGEI demonstrating that this structure has indeed good enough waveguiding properties. Then, in the model described in the present paper, the VLF beam is postulated already on the input of the system. To understand, how such a beam is excited by the, say, dipole antenna near the lower boundary $z=0$ of the WGEI, the formation of the beam structure on the basis of the mode presentation should be searched. Then the conditions of the radiation (absence of ingoing waves) (12) can be used as the boundary conditions for the VLF beam radiated to the upper ionosphere/magnetosphere. Due to relatively large scale of the inhomogeneity in this region, the complex geometrical optics (Rapoport et al. 2014) would be quite suitable for the modeling a beam propagation, even accounting for the wave dispersion in magnetized plasma. The proper effective boundary condition, similarly to (Rapoport et al. 2014) would allow to make relatively accurate matching between the regions, described by means of full wave electromagnetic approach (with Maxwell equations) and complex geometrical optics (FWEM-CGO approach). All of these material is not included into the present paper, but will be delivered in the two future papers, one of which will be dedicated, shortly speaking, to the mode presentation of the VLF propagation in the WGEI, and the other one - to the leakage of VLF beams from the WGEI into
upper ionosphere and magnetosphere and the propagation in these media. But we should mention in the present paper only one result which concerns the mode excitations in the WGEI, because this result is principally important for the justification of TIMEB method. Namely, it is shown that the $\geq 5$ lowest modes of the WGEI are strongly localized in the atmosphere-lower ionosphere. Their longitudinal wavenumbers are close to the corresponding wavenumbers of EM waves in the atmosphere. This fact convinces that the TIMEB method can be applied to the propagation of VLF electromagnetic waves in the WGEI.

4. The Results of Modeling

The dependencies of the components of the permittivity $\varepsilon_x$, $\varepsilon_y$, $\varepsilon_z$ in the coordinate frame associated with the geomagnetic field $\vec{H}_0$ are given in Fig. 3.2. The typical results of simulations are presented in Fig. 4.3. The parameters of the ionosphere correspond to Fig. 3.2. The inclination of the geomagnetic field is $45^\circ$. The VLF frequency is $\omega = 10^5 \text{s}^{-1}$, $f = \omega/2\pi \approx 15.9 \text{kHz}$. The Earth’s surface is assumed as ideally conductive here: $Z = 0$. The values of EM field are given in absolute units, i.e. the magnetic field is measured in Oersteds (Oe), or Gauss (Gs), $1 \text{Gs} = 10^{-4} \text{T}$, whereas the electric field is also in Gs, $1 \text{Gs} = 300 \text{V/cm}$ there.

Note that in the absolute (Gaussian) units the magnitudes of the magnetic field component $|H_y|$ are the same as ones of the electric field component $|E_z|$ in the atmosphere region where the permittivity is $\varepsilon \approx 1$. Below in the figure captions the correspondence between the absolute units and practical SI ones is given.

It is seen that the absolute values of the components of the permittivity increase sharply just above $z = 75 \text{km}$. The behavior of the components of the permittivity is step-like, as seen from Fig. 3.2, a. Due to this, the results of simulations are tolerant to the choice of the position of the upper wall of the waveguide the Earth’s surface – ionosphere. The computed components of the tensor impedance at $z = 85 \text{km}$ are: $Z_{11} = 0.087 + i0.097$, $Z_{21} = 0.085 + i0.063$, $Z_{12} = -0.083 - i0.094$, $Z_{22} = 0.093 + i0.98$. So, a condition $|Z_{ab}| \leq 0.15$ is satisfied there, which is necessary for applicability of the boundary conditions (3).

The maximum value of the $H_z$ component is $0.1 \text{Oe} = 10^{-4} \text{T}$ in Fig. 3, a) for the initial VLF beam at $x = 0$. This corresponds to the value of $E_z$ component of $0.1 \text{Gs} = 30 \text{V/cm}$. At the distance $x = 1000 \text{km}$ the magnitudes of the magnetic field $H_y$ are of about $3\times10^{-4} \text{Oe} = 3 \text{nT}$, whereas ones of the electric field $E_y$ are of about $3\times10^6 \text{Gs} = 1 \text{mV/cm}$.
Figure 3. Part a) is the vertical dependencies of the components of modules of components of the permittivity in the frame associated with the geomagnetic field $|\varepsilon_1|$, $|\varepsilon_3|$, $|\varepsilon_h|$, the curves 1, 2, 3 correspondingly. Parts b) – g) are the real and imaginary parts of the components $\varepsilon_1$, $\varepsilon_3$, $\varepsilon_h$, general and detailed views.

Figure 4. Part a) is the initial distribution of $|H_y|$ at $x = 0$. Parts b), c) are $|E_y|$ and $|H_y|$ at $x = 600$ km. Parts d), e) are $|E_y|$ and $|H_y|$ at $x = 1000$ km. For the electric field it is $3 \times 10^{-6}$ Gs $\approx 1$ mV/cm, for the magnetic field it is $3 \times 10^{-5}$ Gs $\approx 3$ nT. At the altitudes $z < 75$ km it is $|E_z| \approx |H_y|$, so $3 \times 10^{-5}$ Gs $\approx 10$ mV/cm there.

It is seen from Fig. 4, b) - e), that the wave beams are localized within the waveguide the Earth’s surface – ionosphere $0 < z < 75$ km mainly in the regions with the isotropic permittivity. The mutual transformations of the beams of different
polarizations occur near the upper boundary, due to the anisotropy of the ionosphere within the thin layer $75 \text{ km} < z < 85 \text{ km}$, Fig. 4, b), d). These transformations depend on the values of the components of the permittivity of the ionosphere at the altitudes $z > 80 \text{ km}$ and the components of the tensor impedance, so the measurements of the phase and amplitude modulations of different EM components near the Earth’s surface can yield the information on the properties of the lower and even middle ionosphere.

The qualitative effect is changing the polarization, i.e. an occurrence of $E_y$ component of the electric field at small altitudes $z \sim 5 – 10 \text{ km}$ near the Earth’s surface, due to these mutual transformations of EM beams, Fig. 4, parts b), d). Note that the case of the ideal conductivity of the Earth’s surface is considered here, $Z = 0$, so at $z = 0$ the component is $E_y = 0$. If this impedance is $Z \neq 0$, then $E_y$ component occurs also at the Earth’s surface.

The magnitudes of $E_y$ component depend essentially on the values of the electron concentration at the altitudes $z = 75 - 100 \text{ km}$. In Fig. 5, parts a), b), there are different dependencies of the electron concentration $n(z)$, three curves, solid (1), dash (2), and dot (3) ones. The corresponding dependencies of the absolute values of the components of the permittivity are in Fig. 4, parts c), d).

Figure 5. Different profiles of the electron concentrations $n$ used in simulations. The solid, dash, and dot curves correspond to these different profiles. Part a) is the detailed view, b) is a general view. The corresponding profiles of the modules of the components of the permittivity $|\varepsilon_3|$ and $|\varepsilon_h|$ are given in parts c), d).

The distributions of $|E_y|$, $|H_y|$ on $z$, $y$ at $x = 1000 \text{ km}$ are given in Fig. 6. Parts a), b) correspond to the solid (1) curve $n(z)$ in Fig. 5; parts c), d) are for the dash (2) curve; parts e), f) are for the dot (3) curve in Fig. 5. The initial beams of $H_y$ are the same and are given in Fig. 4, a). The values of the tensor impedance for these 3 cases are presented in Table 1.

Table 1. The values of the tensor impedance corresponding to the data shown in Fig. 5. Impedances presented in the lines 1,2 and 3 in Table one correspond to the solid (1), dash (2) and dot (3) curves in Figs. 5 a)-d), respectively.

<table>
<thead>
<tr>
<th>$Z_{11}$</th>
<th>$Z_{12}$</th>
<th>$Z_{21}$</th>
<th>$Z_{22}$</th>
</tr>
</thead>
<tbody>
<tr>
<td>0.088 + i0.098</td>
<td>0.085 + i0.063</td>
<td>-0.083 - i0.094</td>
<td>0.093 + i0.098</td>
</tr>
<tr>
<td>0.114 + i0.127</td>
<td>0.107 + i0.079</td>
<td>-0.105 - i0.127</td>
<td>0.125 + i0.125</td>
</tr>
<tr>
<td>0.067 + i0.0715</td>
<td>0.061 + i0.051</td>
<td>-0.060 - i0.070</td>
<td>0.069 + i0.072</td>
</tr>
</tbody>
</table>
Figure 6. Parts a), c), e) are dependencies of \(|E_y|\), parts b), d), f) are dependencies of \(|H_y|\) at \(x = 1000\) km. The initial beams are the same as in Fig. 4, a). Parts a), b) corresponds to the solid (1) curves in Fig. 5; parts c), d) are for the dash (2) curves; parts e), f) correspond to the dot (3) curves there. For the electric field it is \(3 \cdot 10^{-6} \text{ Gs} \approx 1\) mV/cm, for the magnetic field it is \(3 \cdot 10^{-5} \text{ Gs} \approx 3\) nT. At the altitudes \(z < 75\) km it is \(|E_z| \approx |H_y|\), so \(3 \cdot 10^{-5} \text{ Gs} \approx 10\) mV/cm there.

Figure 7. The dependencies of EM components on the altitude \(z\) in the center of the waveguide \(y = 1500\) km for the different profiles of the electron concentration. The solid (1), dash (2), and dot (3) curves correspond to the different profiles of the electron concentration in Fig. 5, a), b), the same kinds of curves. For the electric field it is \(3 \cdot 10^{-6} \text{ Gs} \approx 1\) mV/cm, for the magnetic field it is \(3 \cdot 10^{-5} \text{ Gs} \approx 3\) nT. At the altitudes \(z < 75\) km it is \(|E_y| \approx |H_y|\), so \(3 \cdot 10^{-5} \text{ Gs} \approx 10\) mV/cm there.

The distributions of \(|E_y|, |H_y|\) on \(z\) at \(x = 1000\) km in the center of the waveguide \(y = 1500\) km are given in Fig. 7. This and other (not presented here) simulations show that change in complex tensors of both volume dielectric permittivity and impedances at the lower and upper boundaries of effective WGEI influence remarkably on the VLF losses in the WGEI. The modulation of the electron concentration at the altitudes above \(z = 120\) km affects weakly the excitation of \(E_y\) component within the waveguide.
5. Discussion

The observations presented in (Roznoi et al. 2015) demonstrated a possibility for seismogenic increasing losses of VLF waves in the WGEI (Fig. 8; see the details in (Roznoi et al. 2015)). We will discuss the correspondence to these experimental results qualitatively.

Figure 8. Averaged through night residual VLF/LF signals in the ground observation for the wave paths: JJY-Moshiri, JJI-Kamchatka, JJY-Kamchatka, NWC-Kamchatka, and NPM-Kamchatka. Horizontal dotted lines show the 2σ level. The color filled zones highlight values exceeding the -2σ level. Two panels below are Dst variations and earthquakes magnitude values (from Roznoi et al., 2015, but not including the DEMETER data). See other details in (Roznoi et al. 2015).

Figure 9. Modification of the ionosphere by electric field of seismogenic origin. (a) – Geometry of the electrostatic problem on the penetration into the ionosphere of the electric field excited by near-ground seismogenic current source; I and II - isotropic and anisotropic regions of the system “atmosphere-ionosphere”. (b) Electric field in the mesosphere; in the presence of the current sources in the mesosphere (curve 1), seismogenic current source in the lower atmosphere (curve 2) and both in the mesosphere and in the lower atmosphere (curve 3); current sources in the mesosphere and lower atmosphere are of the same sign and coincide by the signs with fair weather current (directed vertically downward). (c) Relative perturbations (normalized on the...
corresponding steady-state values in the absence of perturbing electric field, denoted by the index “0”) of electron temperature ($T_e / T_{e0}$), electron concentration ($N_e / N_{e0}$), and electron collision frequency ($\nu_e / \nu_{e0}$).

Figure 10. Altitude distributions of the normalized tangential ($y$) electric (a) and magnetic (b) VLF beam field components in the central plane of the transverse beam distribution ($\phi=0$) at the distance $z=1000$ km from the input of the system. Curves 1 in Fig. a, b correspond qualitatively to the presence of only mesospheric electric current source (with relatively smaller value of $N_e$ and larger $\nu_e$) and curves 2 – to the presence of both mesospheric and near-ground seismogenic electric current sources (with relatively larger value of $N_e$ and smaller $\nu_e$); curves 1 and 2 correspond to the identical input beams at $z=0$ (not shown here).

Curves 1 and 2 in Fig. 10 a, b correspond qualitatively to the curves 1 and 3 in Fig. 9 b, respectively (see also Fig. 9 c and the caption to that figure).

To do this, account for the modification of the ionosphere due to electric field excited by the near-ground seismogenic current source. In the model (Rapoport et al. 2006), the presence of the mesospheric current source, which reflects the observations (Martynenko 2004; Meek 2001; Bragin 1974) is taken into account, and curve 1 in Fig. 9 b the corresponding vertical field distribution in the mesosphere. It is supposed that the mesospheric current has only $Z$ component and is positive, what means that it is directed vertically downward, as well as fair-weather current (curve 1, Fig. 8). Then suppose that near-ground seismogenic current is directed in the same way, as mesospheric current. If the mesospheric current is equal to zero and only corresponding seismogenic near-ground current is present. Corresponding mesospheric electric field, under the condition of given difference of the potentials between the Earth and the ionosphere (curve 2, Fig. 9 b), is directed oppositely to those excited by the corresponding mesospheric current (curve 1, Fig. 9 b). As a result, in the presence of both mesospheric and seismogenic near-ground current, the total mesospheric electric field (curve 3, Fig. 8 b) is less by the absolute value, than those in the presence of only mesospheric current (curve 1, Fig. 9 b). As it is shown in (Rapoport et al. 2006), the decrement of losses $|k^*|$ for VLF waves in the WGEI is proportional to $|k^*| \cdot |\nu_e| \cdot |N_e| / \nu_e$. Accounting for that in the external electric field in the mesosphere $N_e$ and $\nu_e$ decreases and increases, respectively, due to the appearance of seismogenic near-ground electric current, in addition to the mesospheric current (curve 3, Fig. 9 b), losses increases comparatively to the case, when the seismogenic current is absent and electric field is larger by absolute value (curve 1, Fig. 9 b). An increasing in VLF beam losses, shown in Fig 10 corresponds to increasing losses with increasing absolute value of imaginary part of the dielectric permittivity when near-ground seismogenic current source (curves 2 in Fig. 9 a, b) appear, additionally to the pre-existing mesospheric current source (curves 1 in Fig. 9 a, b, see also caption to the Fig. 9). This corresponds qualitatively to the results presented in (Roznoi et al. 2015), see also Figs. 8. Note that the above mentioned
estimations concern only volume losses in the WGEI. Losses connected with the modification of effective impedance are not included in the course of these elementary estimations, and the more detailed consideration of both the modification of the ionosphere by means of electric-photochemistry mechanism and the variation of losses due to all mechanisms including volume and effective impedance effects will be a subject of the subsequent papers.

The closest approach to a direct investigation of the profile of VLF electromagnetic field in the Earth-Ionosphere waveguide was a series of sounding rocket campaigns at mid- and high-latitudes [Wallops Is., VA, and Siple Station, Antarctica - Kintner et al., 1983; Brittain et al., 1983; Siefring and Kelly, 1991; Arnoldy and Kintner, 1989] where single-axis E-field and three-axis B-field antennas, supplemented in some cases with in situ plasma density measurements were used to detect the far-field fixed-frequency VLF signals radiated by US Navy and Stanford ground transmitters.

The most comprehensive study of the WGEI will be provided by the ongoing NASA VIPER (VLF Trans-Ionospheric Propagation Experiment Rocket) project (PI J. W. Bonnell, UC Berkeley, NASA Grant 80NSSC18K0782). The VIPER sounding rocket campaign is consist of a summer nighttime launch during quiet magnetosphere conditions from Wallops Flight Facility, VA, collecting data through the D, E, and F regions of the ionosphere with a payload carrying the following instrumentation: 2D E- and 3D B-field waveforms, DC-1 kHz; 3D ELF to VLF waveforms, 100 Hz to 50 kHz; 1D wideband E-field measurement of plasma and upper hybrid lines, 100 kHz to 4 MHz; and Langmuir probe plasma density and ion gauge neutral density measurements at a sampling rate of at least tens of Hz. The VIPER project will fly a fully 3D EM field measurement, DC through VLF, and relevant plasma and neutral particle measurements at mid-latitudes through the radiation fields of (1) an existing VLF transmitter (the VLF transmitter Cutler with call sign NAA, the very low frequency (VLF) shore radio station at Cutler, Maine, USA, which transmits, at a frequency of 24 kHz an input power of up to 1.8 megawatts, see Fig. 11) and (2) naturally-occurring lightning transients through and above the leaky upper boundary of the WGEI supported by a vigorous theory and modelling effort in order to explore the vertical and horizontal profile of the observed 3D electric and magnetic radiated fields of the VLF transmitter, and the profile related to the observed plasma and neutral densities. The VLF waves reflection, absorption, and transmission processes as a function of altitude will be searched making use of the data on the vertical VLF E- and B-field profile.
Figure 11. Proposed VIPER Trajectory

The aim of this experiment is the investigation of the VLF beams launched by the near-ground source/VLF transmitter with the known parameters and propagating both in the WGEI and leaking from WGEI into the upper ionosphere. Characteristics of these beams will be compared with the theory proposed in the present paper and the theory on leakage of the VLF beams from WGEI, which we will present in the next papers.

Conclusions

1. We have developed the new and highly effective robust method of tensor impedance for the VLF electromagnetic beam propagation in the inhomogeneous waveguiding media - ’’tensor impedance method for modelling propagation of electromagnetic beams (TIMEB)’’ in a multi-layered/inhomogeneous waveguide.

2. The main differences/advantages of the proposed tensor impedance method in comparison with the known method of the impedance recalculating, in particular invariant imbedding methods (Shalashov and Gospodchikov, 2010; Kim and Kim, 2016) are the following: (i) our method is a direct method of the recalculation of tensor impedance, and the corresponding tensor impedance is determined analytically, see Eqs. (22). (ii) Our method, for the media without non-locality, does not need a solution of integral equation(s), as the invariant imbedding method does (iii) The proposed tensor impedance method does not need the revealing of forward and reflected waves. Moreover, even the conditions of the radiation (12) at the upper boundary $z=L_{max}$ is determined through the total field components $H_{x,y}$, what makes the proposed procedure technically less cumbersome and practically more convenient.

3. The application of this method jointly with the previous results of the modification of the ionosphere by seismogenic electric field gives the results, which qualitatively are in an agreement with the experimental data on the seismogenic increasing losses of VLF waves/beams propagating in the WGEI.

4. The waveguide includes the region for the altitudes $0 < z < 80 - 90$ km. The boundary conditions are the radiation conditions at $z = 300$ km, they can be recalculated to the lower altitudes as the tensor relations between the tangential components of the EM field. In another words, the tensor impedance conditions have been used at $z = 80 – 90$ km.
The observable qualitative effect is mutual transformations of different polarizations of the electromagnetic field occur during the propagation. This transformation of the polarizations depends on the electron concentration, i.e. the conductivity, of D- and E-layers of the ionosphere at the altitudes 75 – 120 km.

Change in complex tensors of both volume dielectric permittivity and impedances at the lower and upper boundaries of effective WGEI influence remarkably on the VLF losses in the WGEI.

The proposed model and results on propagation of VLF electromagnetic beams in the WGEI will be useful to explore the characteristics of these waves as an effective instrument for diagnostics the influences on the ionosphere “from above” (Sun-Solar Wind-Magnetosphere-Ionosphere), “from below” (the most powerful meteorological, seismogenic and other sources in the lower atmosphere and lithosphere/Earth, such as hurricanes, earthquakes, tsunamis etc.), from inside the ionosphere (strong thunderstorms and lightning discharges) and even from the far space (such as gamma-flashes, cosmic rays etc.).

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REFERENCES


Appendix: the matrix coefficients included into eq. (16)

Here the expressions of the matrix coefficients are presented that are used in the matrix factorization to compute the tensor impedance, see eq. (16).

\[
\hat{\alpha}'(m) = \begin{pmatrix}
\frac{i k h_{1}}{\Delta}(k_{1} - \alpha_{1} k_{2}) + \frac{i h_{1}}{\Delta} a_{1}(k_{2} - k_{1}) \\
\frac{i h_{1}}{\Delta} a_{1}(k_{1} - k_{2}) + 1 + \frac{i h_{1}}{\Delta} (k_{2} - \alpha_{2} k_{1})
\end{pmatrix}, \quad \Delta = 1 - \alpha_{1} \alpha_{2};
\]

\[
\hat{\alpha}'(+) = \begin{pmatrix}
-1; 0 \\
0; -1
\end{pmatrix}
\]

\[
\hat{\alpha}'(+) = \begin{pmatrix}
\left(1 - \beta_{21} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} \\
\left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}}
\end{pmatrix}
\]

\[
\hat{\alpha}'(+) = \begin{pmatrix}
\left(1 - \beta_{21} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} \\
\left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}}
\end{pmatrix}
\]

\[
\hat{\alpha}'(+) = \begin{pmatrix}
\left(1 - \beta_{21} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} \\
\left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}}
\end{pmatrix}
\]

\[
\hat{\alpha}'(+) = \begin{pmatrix}
\left(1 - \beta_{21} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} \\
\left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}}
\end{pmatrix}
\]

\[
\hat{\alpha}'(+) = \begin{pmatrix}
\left(1 - \beta_{21} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} \\
\left(-\beta_{21} \right)_{j_{1},j_{2}} + \frac{i k h_{1}}{\Delta 2} \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}} & \left(1 - \beta_{22} k_{1}^{2} k_{0}^{-2}\right)_{j_{1},j_{2}}
\end{pmatrix}
\]
\[ \hat{\alpha}_j^{(1)} = \begin{pmatrix} \frac{\beta_{22}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} & -\left( \frac{\beta_{22}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \frac{i k_j h_j}{2} & \left( \frac{\beta_{23}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \\ \frac{\beta_{32}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} & \frac{\beta_{32}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \frac{i k_j h_j}{2} & \left( \frac{\beta_{33}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \\ \frac{\beta_{11}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} & \frac{\beta_{11}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \frac{i k_j h_j}{2} & \left( \frac{\beta_{12} \beta_{23}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \end{pmatrix} \]

\[ \hat{\alpha}_j^{(0)} = \begin{pmatrix} \frac{\beta_{22}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} - \left( \frac{\beta_{22}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j-1/2} + \frac{k_j^2 h_j^2}{k_0^2} & \left( \frac{\beta_{23}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j-1/2} + \left( \frac{\beta_{23}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \\ \frac{\beta_{32}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} + \left( \frac{\beta_{32}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \frac{k_j^2 h_j^2}{k_0^2} & \frac{\beta_{33}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} - \left( \frac{\beta_{33}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j-1/2} - \left( \frac{\beta_{33}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \\ \frac{\beta_{11}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} - \left( \frac{\beta_{11}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j-1/2} - \left( \frac{\beta_{11}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} + \frac{k_j^2 h_j^2}{k_0^2} \left( \frac{\beta_{12} \beta_{23}}{1-\beta_{22} \frac{k_j^2}{k_0^2}} \right)_{j+1/2} \end{pmatrix} \]